













Quasi-Fermi Energies
$$E_F^h = E_V + kT \cdot \ln \frac{N_{eff}^h}{n^{h,*}}$$
 $E_F = E_V + kT \cdot \ln \frac{N_{eff}^h}{n^h}$  $E_F - E_F^h = kT \cdot \ln \frac{n^{h,*}}{n^h}$ 

























## Gaertner's Model

Furthermore, if  $\alpha^{-1}$  << W (weakly absorbed light), then

$$j_{photo} = -q\Phi(1 - \frac{1}{1 + \alpha Lp}(1 - \alpha W)) =$$
$$= -q\alpha\Phi\frac{Lp + W}{1 + \alpha Lp}$$

If further Lp<<W (high recombination in the quasi-neutral region or low mobility), then

$$j_{photo} = -q \alpha \Phi W$$

The photocurrent is proportional to the width of the space-charge layer! 21

## Gaertner's Model

$$j_{photo} = -q \alpha \Phi W$$

We remember from earlier lectures that;

$$L_{sc} \equiv \Phi = \sqrt{\frac{2\varepsilon\varepsilon_0}{e_0 N_D}} \left| \phi - \phi_{fb} \right|$$

The photocurrent is proportional to the square root of the bias applied across the space-charge region and will also increase with a decrease in the donor density (wider SCR).

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